

A geometric framework for interpreting and parameterising geostrophic ocean eddies

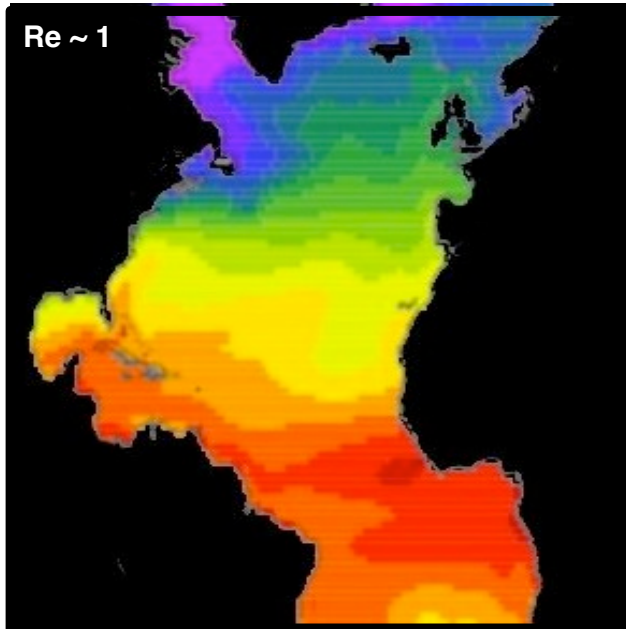


David Marshall (*University of Oxford*),

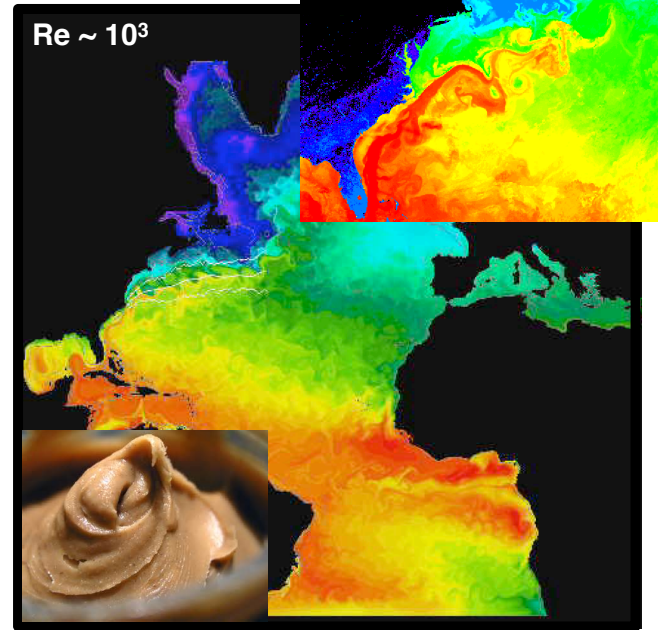
James Maddison (*University of Edinburgh*)



1^0 (climate) resolution



$1/12^0$ resolution



(MICOM, University of Miami)

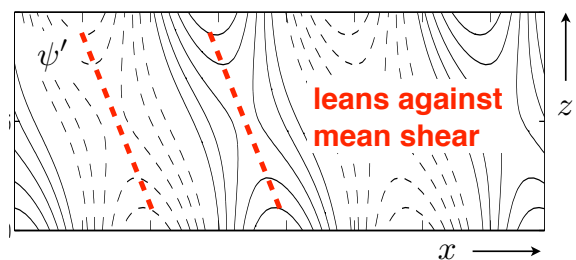
Classical paradigm for location/structure of ocean eddies:

Eady (1949) model of baroclinic instability

- uniform rotation
- uniform stratification
- uniform shear
- opposing potential vorticity gradients at upper and lower boundaries



most **unstable** mode:



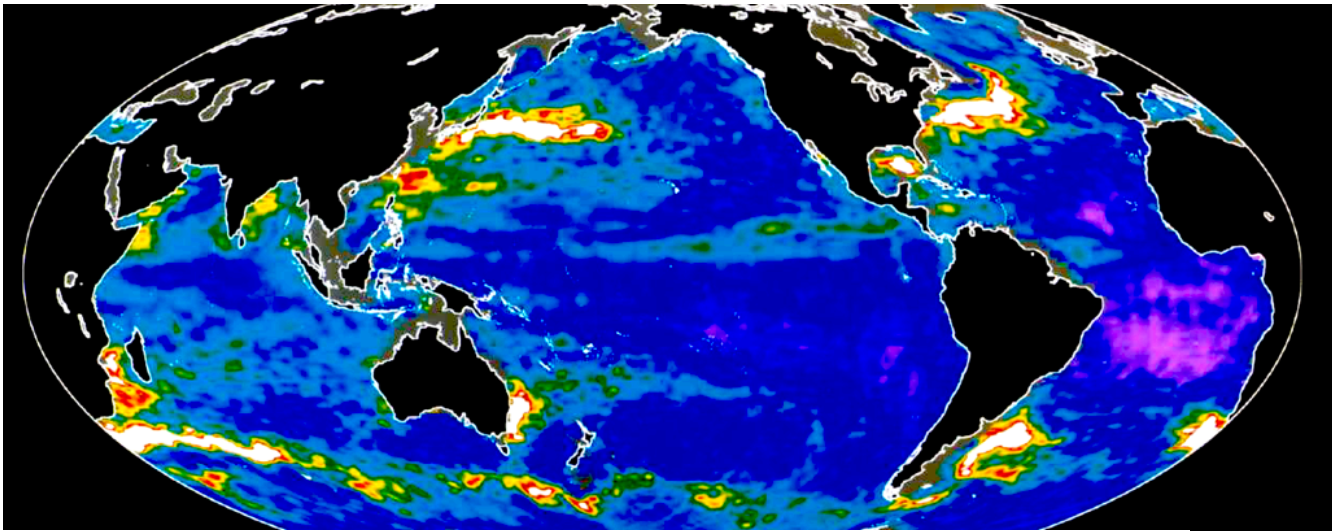
(figure: adapted from Vallis 2006)

energy growth rate for most unstable mode:

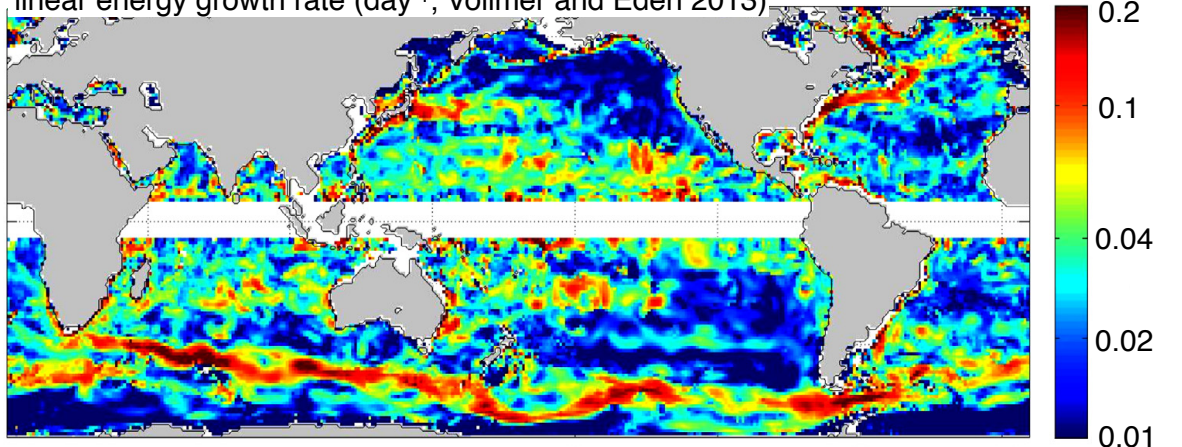
$$0.61 \frac{f_0}{N_0} \frac{\partial u}{\partial z} \sim \begin{matrix} 0.3 \text{ day}^{-1} & \text{- atmosphere} \\ 0.03 \text{ day}^{-1} & \text{- ocean} \end{matrix}$$

length scale of instability characterised by
Rossby deformation radius:

$$L_d = \frac{N_0 H}{f_0} \sim \begin{matrix} 1000 \text{ km} & \text{- atmosphere} \\ 50 \text{ km} & \text{- ocean} \end{matrix}$$



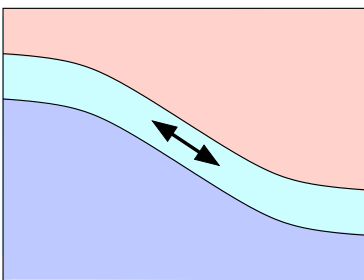
linear energy growth rate (day⁻¹; Vollmer and Eden 2013)



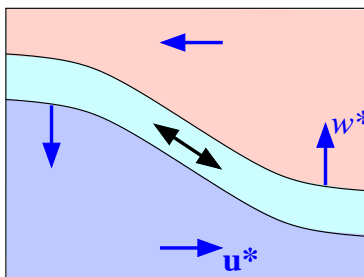
Gent and McWilliams (1990):

adiabatic parameterisation of baroclinic instability

eddies mix along isopycnals (Redi 1982) ...



... and advect by an *eddy bolus velocity* - flattens isopycnals (Gent et al. 1995)



$$\mathbf{u}^* = \frac{\partial}{\partial z} \left(\kappa \frac{\nabla b}{\partial b / \partial z} \right), \quad w^* = -\nabla \cdot \left(\kappa \frac{\nabla b}{\partial b / \partial z} \right),$$

removes available potential energy

can relate eddy diffusivity, κ , to mean flow (e.g., Visbeck et al. 1997)
or eddy energy (Eden and Greatbatch 2008)

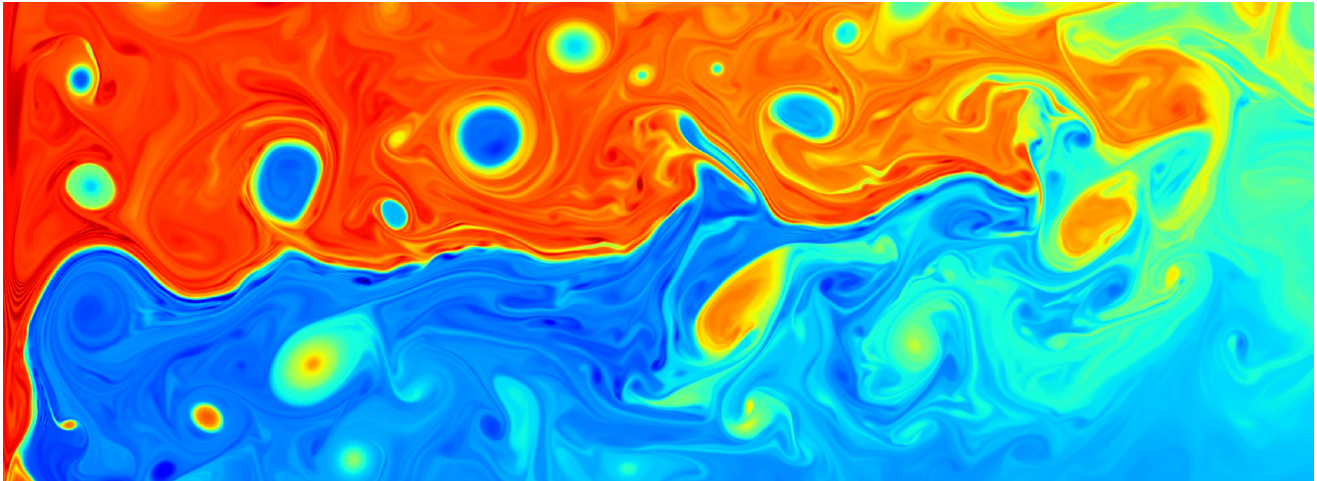
Alternative paradigm: potential vorticity mixing

often advocated ... rarely implemented in ocean GCMs!

Idea: potential vorticity $q = \frac{f + \xi}{h}$ is materially conserved in absence of forcing/dissipation:

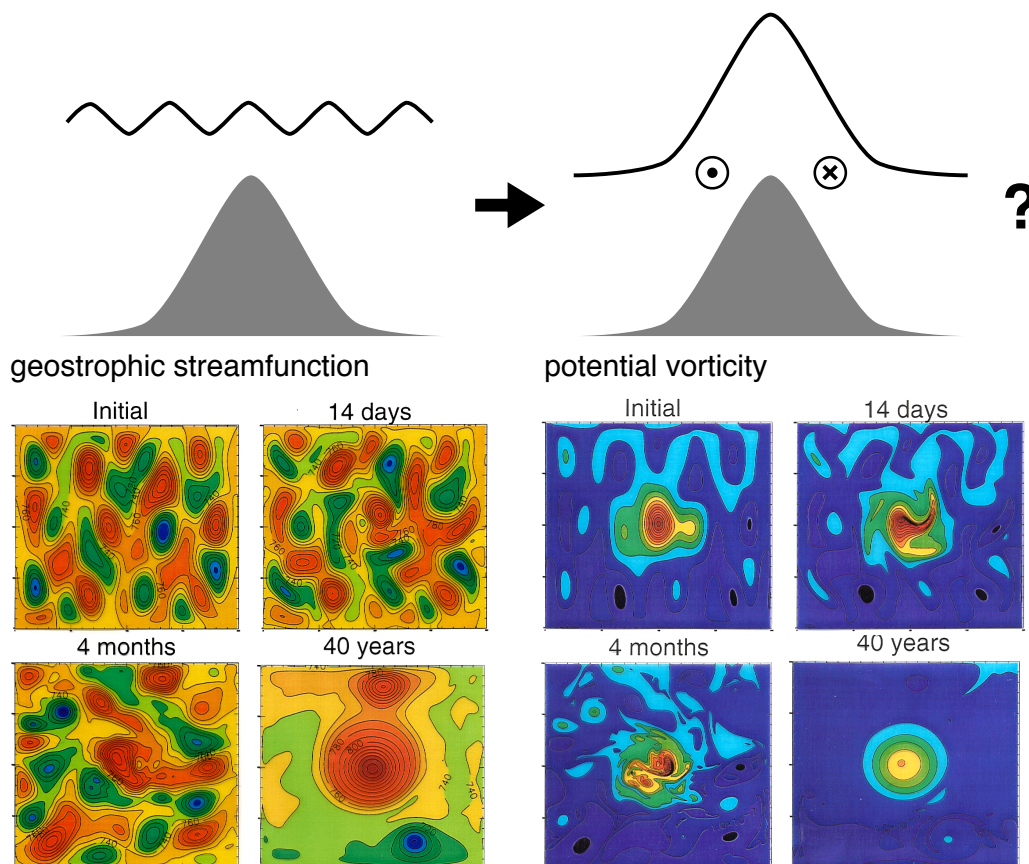
$$\frac{\partial q}{\partial t} + \mathbf{u} \cdot \nabla q = 0$$

stirred and mixed along isopycnals \Rightarrow down-gradient closure, $\overline{q' \mathbf{u}'} = -\kappa \nabla_{\rho} \bar{q}$?



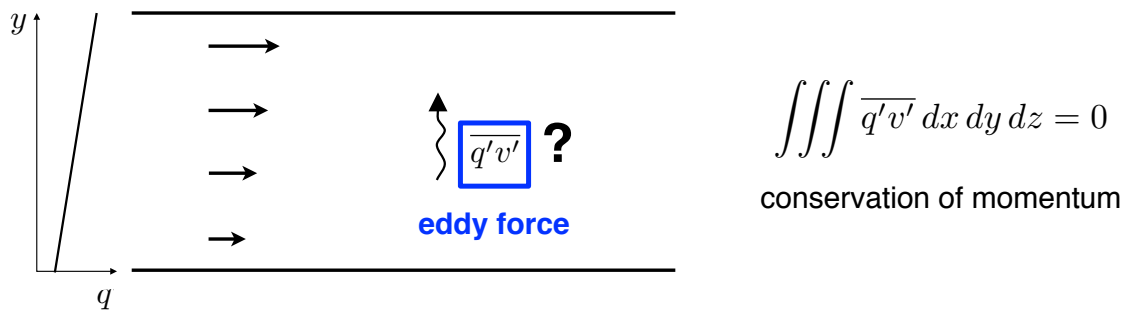
PV mixing problem 1: conservation of energy

e.g. , freely-decaying turbulence over a seamount (Adcock and Marshall, 2000)



PV mixing problem 2: conservation of momentum

periodic channel:



not satisfied by down-gradient potential vorticity closures without additional constraints
(Green, 1970; J. Marshall, 1981)

e.g., here $\overline{q'v'} = -\kappa \partial q / \partial y$ only consistent if $\kappa = 0$

note: this is the Charney-Stern stability condition

Take-home message:

Eddies mix potential vorticity along density surfaces ...
... subject to constraints of energy and momentum conservation

Goal of this work:

Develop framework for **interpreting** and **parameterising** eddy potential vorticity fluxes
in which the relevant **symmetries and conservation laws are preserved**.

Work with quasi-geostrophic “residual-mean” equations:

$$\frac{\partial \bar{\mathbf{u}}_g}{\partial t} + \dots = \boxed{-\mathbf{k} \times \overline{q' \mathbf{u}'}}$$

eddy force

how to parameterise?

(Maddison and Marshall, 2013; cf. Young, 2012)

Key idea:

Write **eddy potential vorticity flux** (or eddy force)
as **divergence of an eddy stress tensor**:

$$\overline{q'\mathbf{u}'} = \nabla \cdot \begin{pmatrix} -N & M - P & 0 \\ M - P & N & 0 \\ R & S & 0 \end{pmatrix} \quad (\text{Plumb 1986})$$

“Taylor identity”

where: $M = \frac{\overline{v'^2 - u'^2}}{2}$ $N = -\overline{u'v'}$ Reynolds stresses

$$P = \frac{\overline{b'^2}}{2\mathcal{N}_0^2} \quad \text{eddy potential energy}$$

$$R = \frac{f_0}{\mathcal{N}_0^2} \overline{u'b'} \quad S = \frac{f_0}{\mathcal{N}_0^2} \overline{v'b'} \quad \text{eddy buoyancy flux / “eddy form stress”}$$

Why do this?!!!

$$\overline{q'\mathbf{u}'} = \nabla \cdot \begin{pmatrix} -N & M - P & 0 \\ M - P & N & 0 \\ R & S & 0 \end{pmatrix}$$

1. This is a mathematical identity! (down-gradient flux \neq divergence of a tensor)

2. Momentum constraints preserved with appropriate boundary conditions:

$$\frac{\partial \overline{\mathbf{u}}_g}{\partial t} + \dots = \nabla \cdot (\text{eddy momentum fluxes})$$

3. Reduces to Gent and McWilliams (1990) / Greatbatch and Lamb (1990)
if we parameterise only the vertical momentum fluxes.

Therefore a natural framework for extending Gent and McWilliams.

Why do this?!!!

$$\overline{q'\mathbf{u}'} = \nabla \cdot \begin{pmatrix} -N & M-P & 0 \\ M-P & N & 0 \\ R & S & 0 \end{pmatrix}$$

4. Suppose we solve an **eddy energy** equation (Eden and Greatbatch, 2008):

$$\frac{\partial \overline{E}}{\partial t} + \dots = -\overline{\mathbf{u}}_g \cdot \nabla (\text{eddy force}) = \overline{\mathbf{u}}_g \cdot \mathbf{k} \times \overline{q'\mathbf{u}'}$$

This **eddy energy** gives a bound on the **magnitude of the eddy stress tensor**:

$$\frac{1}{2} \left[(-N)^2 + (M-P)^2 + (M+P)^2 + N^2 + \frac{\mathcal{N}_0^2}{f_0^2} (R^2 + S^2) \right] \leq E^2$$

This means there are **no remaining dimensional unknowns**!

Why do this?!!!

$$\overline{q'\mathbf{u}'} = \nabla \cdot \begin{pmatrix} -N & M-P & 0 \\ M-P & N & 0 \\ R & S & 0 \end{pmatrix}$$

5. This allows us to rewrite the eddy stress tensor in terms of the **eddy energy**, two non-dimensional **eddy anisotropies**, and three **eddy angles**:

$$M = -\gamma_m E \cos 2\phi_m \cos^2 \lambda \quad N = \gamma_m E \sin 2\phi_m \cos^2 \lambda \quad P = E \sin^2 \lambda$$

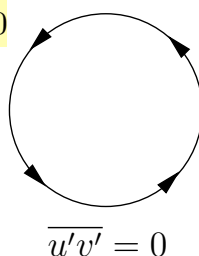
$$R = \gamma_b \frac{f_0}{\mathcal{N}_0} E \cos \phi_b \sin 2\lambda \quad S = \gamma_b \frac{f_0}{\mathcal{N}_0} E \sin \phi_b \sin 2\lambda$$

horizontal angles

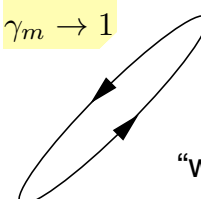
vertical angle

e.g., barotropic eddies:
(plan view)

$$\gamma_m = 0$$



$$\gamma_m \rightarrow 1$$

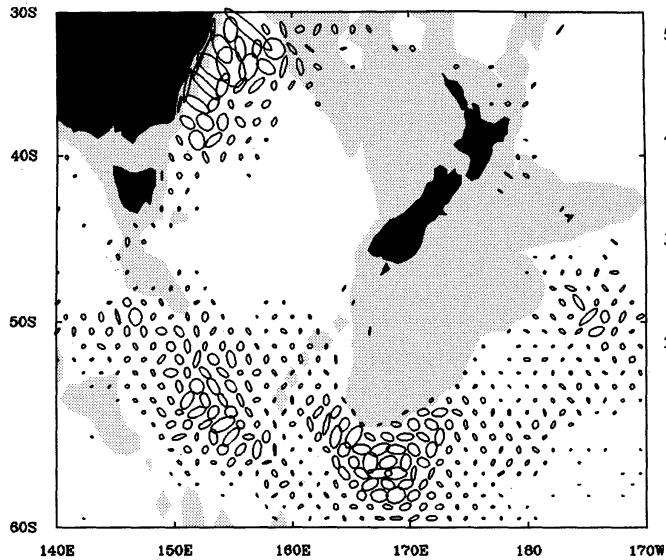


$$\overline{u'v'} > 0$$

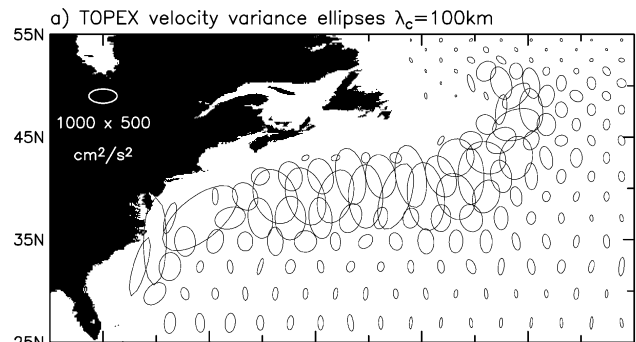
Why do this?!!!

$$\overline{q'\mathbf{u}'} = \nabla \cdot \begin{pmatrix} -N & M-P & 0 \\ M-P & N & 0 \\ R & S & 0 \end{pmatrix}$$

Equivalent to **eddy variance ellipses** used with altimetric / float data:



(Morrow et al. 1994)



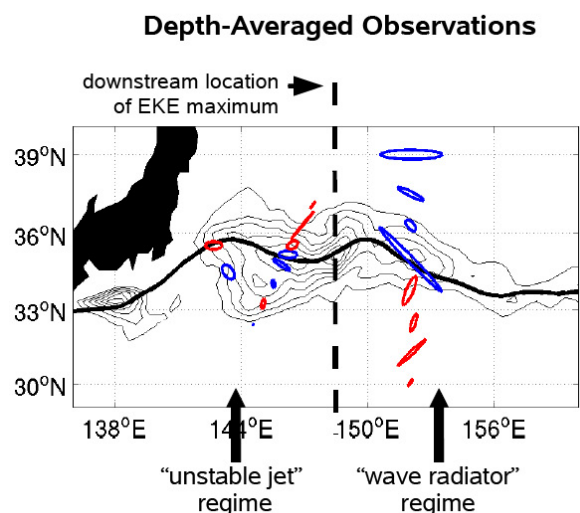
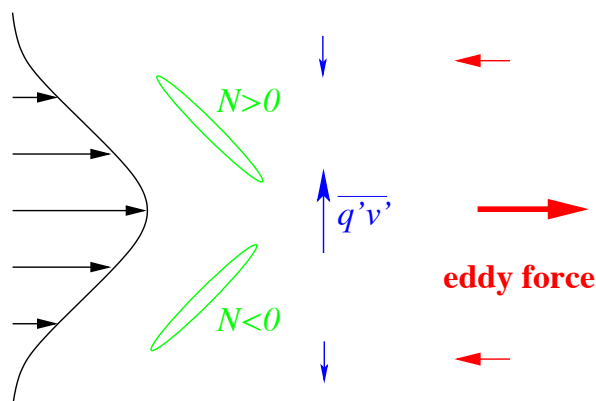
(Chelton and Schlax 2003)

Why do this?!!!

$$\overline{q'\mathbf{u}'} = \nabla \cdot \begin{pmatrix} -N & M-P & 0 \\ M-P & N & 0 \\ R & S & 0 \end{pmatrix}$$

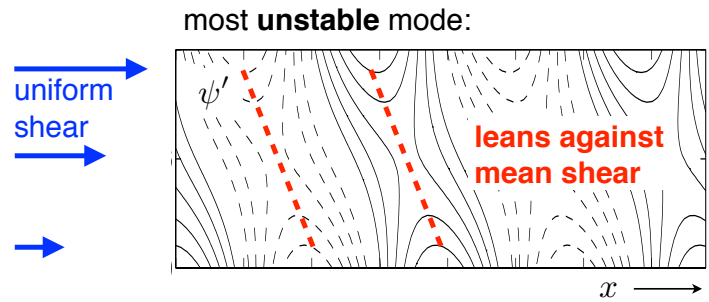
6. Eddy angles have a strong connection with classical stability theory:

eddies lean “against” mean shear \Rightarrow extract energy from mean flow - **instability**;
eddies lean “into” mean shear \Rightarrow return energy to mean flow - **stability**.



(Waterman et al. 2011)

Application: Eady model



Eddy energy budget:

$$\frac{\partial}{\partial t} \iiint E \, dx \, dy \, dz = - \iiint \bar{u} \overline{q'v'} \, dx \, dy \, dz$$

$\mathbf{u} \cdot$ (eddy force)

$$= - \iiint \bar{u} \frac{\partial S}{\partial z} \, dx \, dy \, dz$$

substitute eddy stress tensor

$$= \iiint \frac{\partial \bar{u}}{\partial z} S \, dx \, dy \, dz$$

integrate by parts

$$= \alpha \frac{f_0}{N_0} \frac{\partial \bar{u}}{\partial z} \iiint E \, dx \, dy \, dz$$

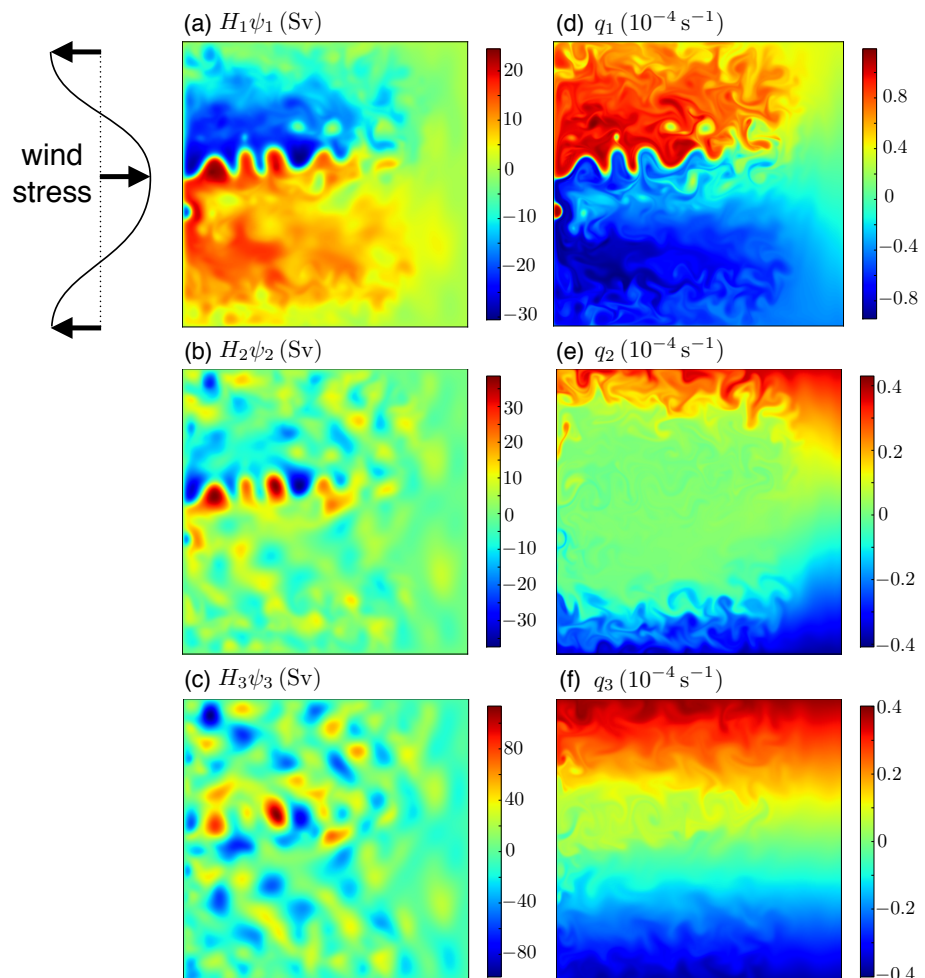
apply energy bound

$$-1 \leq \alpha \leq 1$$

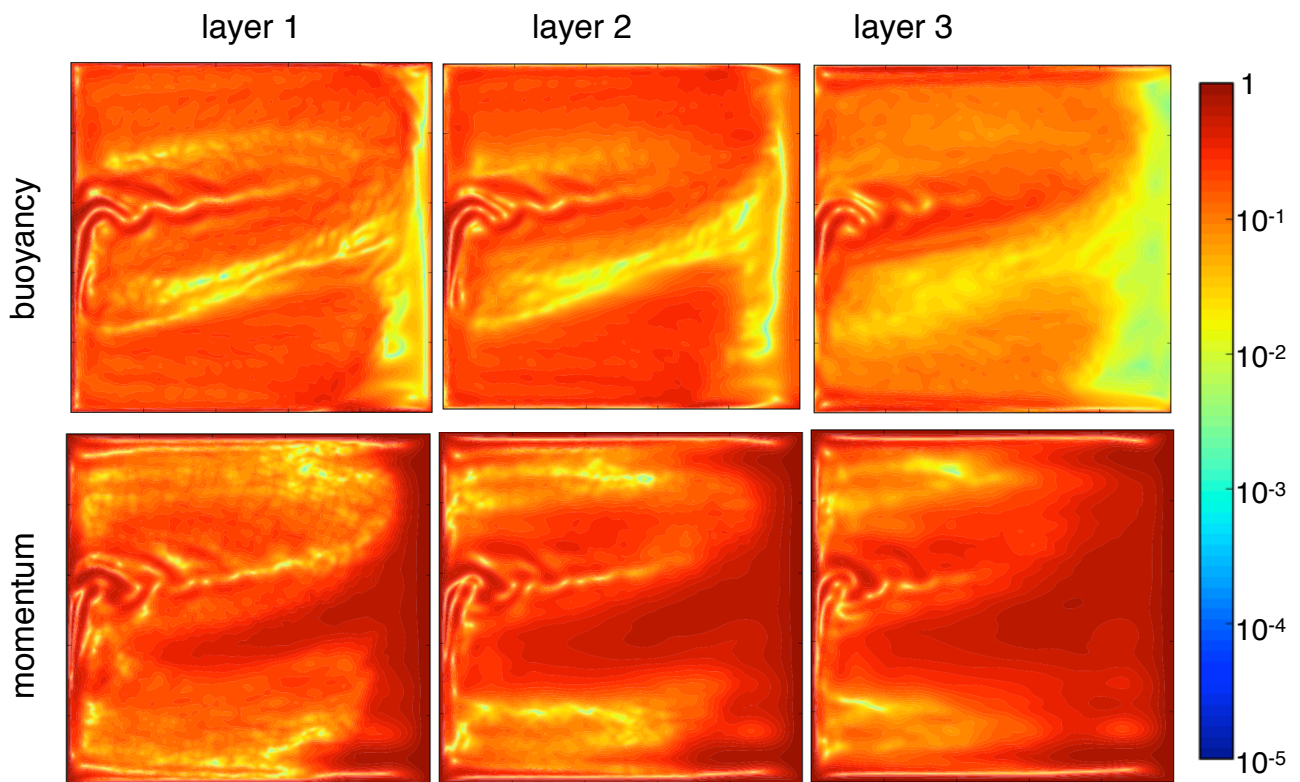
Eady growth rate
if $\alpha = 0.61$

How anisotropic
are the eddies?

3-layer QG model



eddy anisotropies



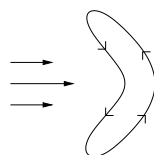
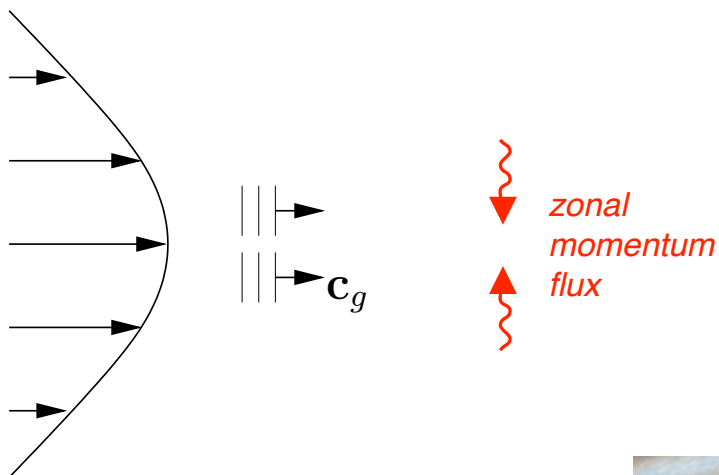
$$N = \gamma_m E \sin 2\phi_m \cos^2 \lambda$$

$$S = \gamma_b \frac{f_0}{N_0} E \sin \phi_b \sin 2\lambda$$

What sets the eddy angles?

- for linear Rossby waves: **refraction**

(e.g., Buhler and McIntyre, 2005)

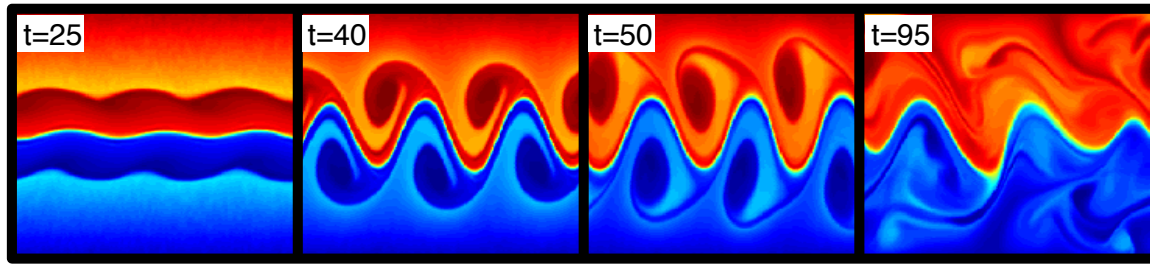


Ray-tracing - barotropic jet

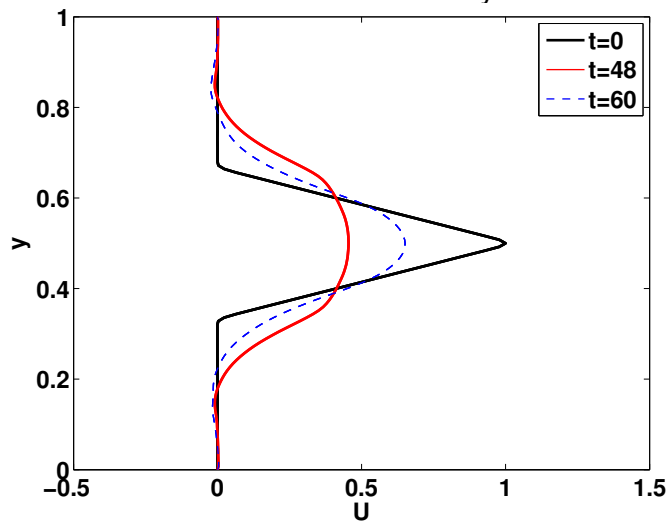
Talia Tamarin



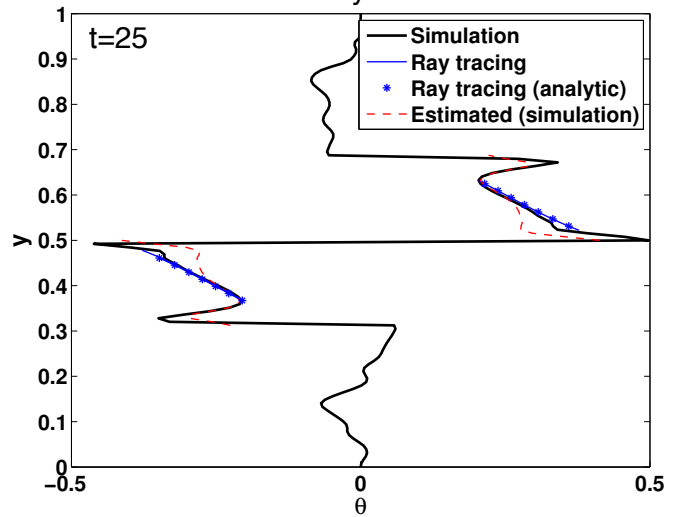
potential vorticity



zonal-mean velocity



eddy tilt

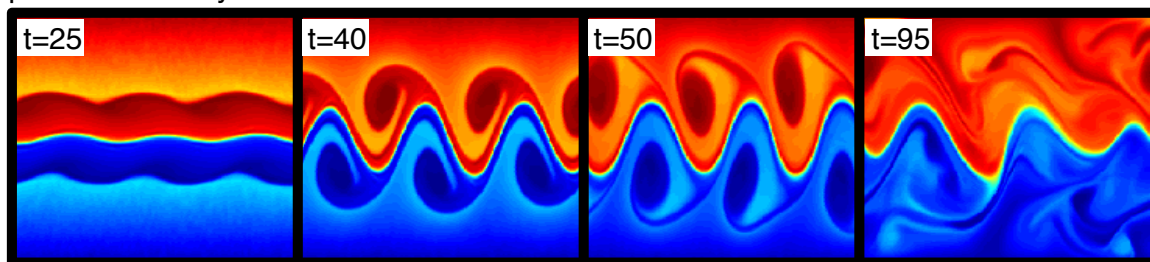


Ray-tracing - barotropic jet

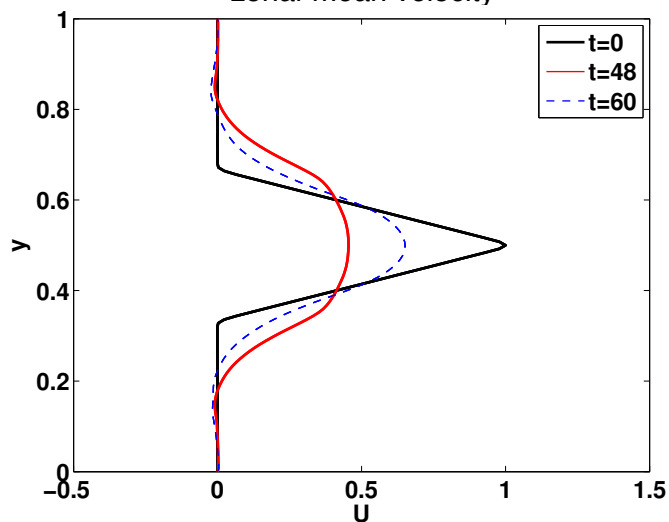
Talia Tamarin



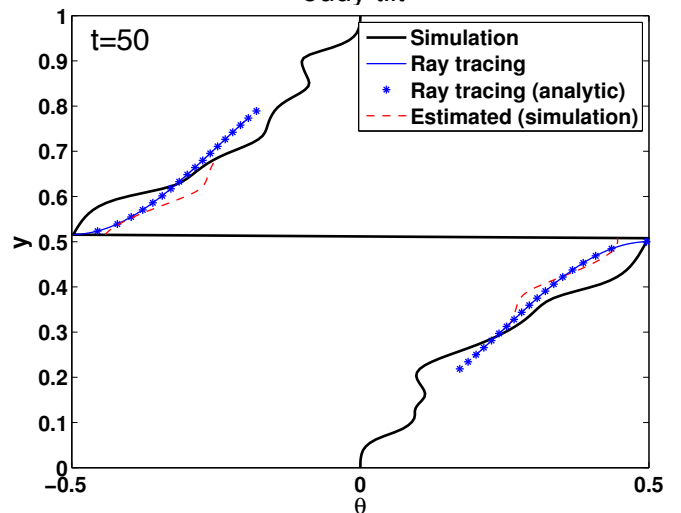
potential vorticity



zonal-mean velocity



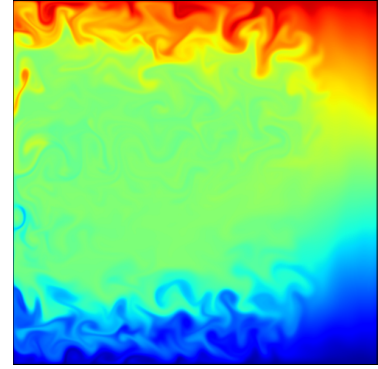
eddy tilt



Mixing of potential vorticity?

- If we: (i) solve an eddy potential enstrophy ($\overline{q'^2}$) budget;
 (ii) include dissipation of $\overline{q'^2}$ (= potential vorticity mixing);
 (ii) ensure $\overline{q'\mathbf{u}'}$ vanishes when $\overline{q'^2}$ vanishes;
 [use another bound on divergence of eddy stress tensor?]

then **Arnold's first stability theorem is preserved.**



Physical interpretation? (Marshall and Adcroft, 2010)

Eddy energy equation:
$$\frac{\partial}{\partial t} \frac{\overline{\mathbf{u}' \cdot \mathbf{u}'}}{2} + \nabla \cdot (\dots) = +\overline{q'\mathbf{u}'} \cdot \nabla \overline{\psi}$$

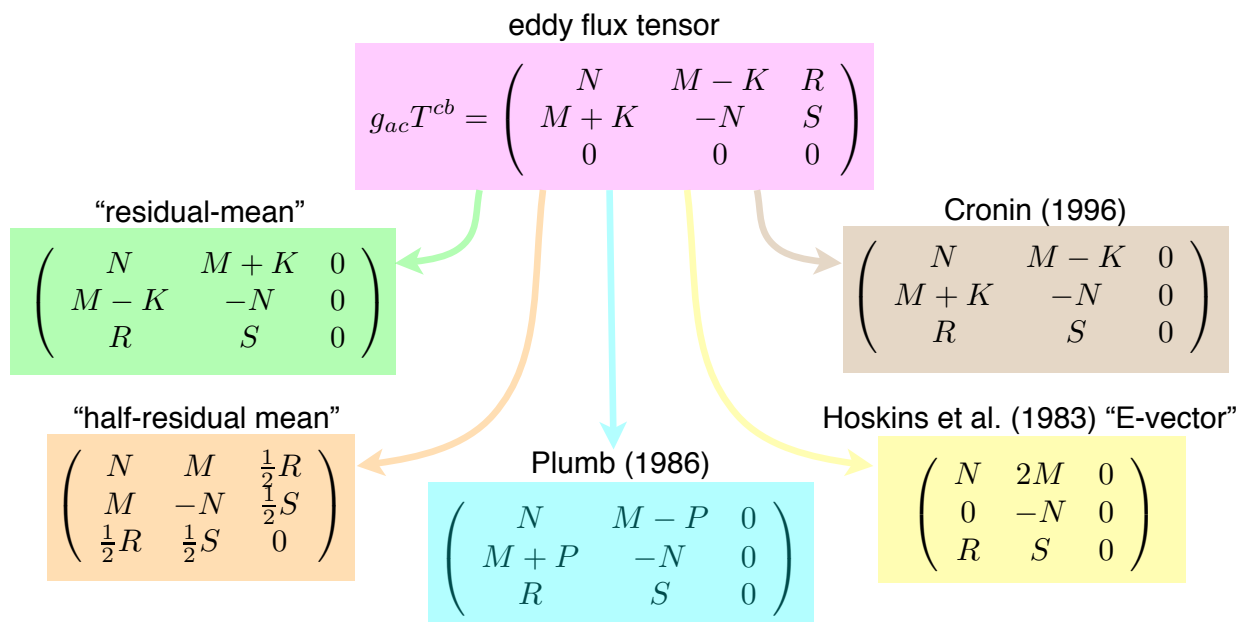
Eddy enstrophy equation:
$$\frac{\partial}{\partial t} \frac{\overline{q'^2}}{2} + \nabla \cdot (\dots) = -\overline{q'\mathbf{u}'} \cdot \nabla \overline{q}$$

If $d\overline{q}/d\overline{\psi} > 0$, eddy energy can grow only at the expense of eddy potential enstrophy.

⇒ stable (in the sense of Lyapunov) - **Arnold's first stability theorem.**

Coordinate-invariant derivation (Maddison and Marshall, 2013)

quasigeostrophic PV equation:
$$\partial_t \overline{q} + \left(\overline{[u_g]^a} \overline{q} \right)_{;a} = -T_{;ab}^{ab}$$
 double divergence
 ⇒ 2 forms of gauge freedom



Approach generalises to isopycnal thickness-weighted primitive equations

(cf. Young, 2012)

Summary of key points

- Geostrophic eddies are fundamental in setting the structure and circulation of the ocean.
- Preserving symmetries and conservation laws in models with parameterised eddies
⇒ classical stability conditions carry over.
- Down-gradient eddy potential vorticity flux closures are inconsistent with the underlying mathematical structure of the eddy-mean flow interaction.
- Gent and McWilliams is consistent with this underlying mathematical structure.
- New geometric framework for diagnosing and interpreting eddy-mean flow interactions.
- Much left to do, e.g.:
 - simple extension of Gent and McWilliams to include up-gradient momentum fluxes;
 - adjoint methods to optimise choice of parameters (Julian Mak);
 - eddy-topography interactions;
 - diagnostics of eddy-mean flow interactions in Southern Ocean (Andreas Klocker);
 - applications to planetary atmospheres?
 - ...

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